

KIDA '17

New calculations of radiative charge transfer and of radiative association rate coefficients

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Outline

Introduction – collision ingredients

Carbon and helium ion (C-He⁺)

Carbon and proton (C-H⁺)

Silicon and oxygen (Si-O)

Introduction

“Recombination with emission of photon”

“Radiative recombination”

Or “**radiative association**” for molecule formation

Atom-atom collision

$A+B \rightarrow AB + \text{photon}$ radiative association

Atom-ion collision

$A+B^+ \rightarrow AB^+ + \text{photon}$ radiative association

$A+B^+ \rightarrow A^++B + \text{photon}$ radiative charge transfer

Ingredients

Scattering treatment

Potential energy curves

Long-range interactions

Transition dipole matrix elements

“oscillator strength”

Wave functions

Atom-ion C-He⁺



For modeling CO formation in SN ejecta, important to know if He⁺ removed by C, S, Si, or O reactions

CO is removed by



Radiative loss cross sections

Essentially no radiative association

Use an optical potential approach for radiative charge transfer:

$$\sigma(E) \sim \frac{\pi}{k^2} \sum_J (2J + 1) [1 - e^{-4\eta_J}]$$
$$\eta_{J(E)} = \frac{\pi}{2} \int_0^\infty dR |F_J(kR)|^2 A(R)$$

Rate coefficients

$$\alpha(T) = \left(\frac{8}{\mu\pi}\right)^{\frac{1}{2}} \left(\frac{1}{k_B T}\right)^{\frac{3}{2}} \int_0^{\infty} dE E \sigma(E) e^{-E/k_B T}$$

Re-Introduction

$C+He^+ \rightarrow C^++He + \text{photon}$ radiative charge transfer

$C+He^+ \rightarrow C^++He$ nonradiative charge transfer

C-He⁺ Kimura *et al.* (PRA 1993, PRA 1994)

O-He⁺ Kimura *et al.* PRA (1994), Zhao et al. (2004)

Si-He⁺ Satta *et al.* MNRAS (2013)

S-He⁺ ...

C-He⁺

Radiative
charge

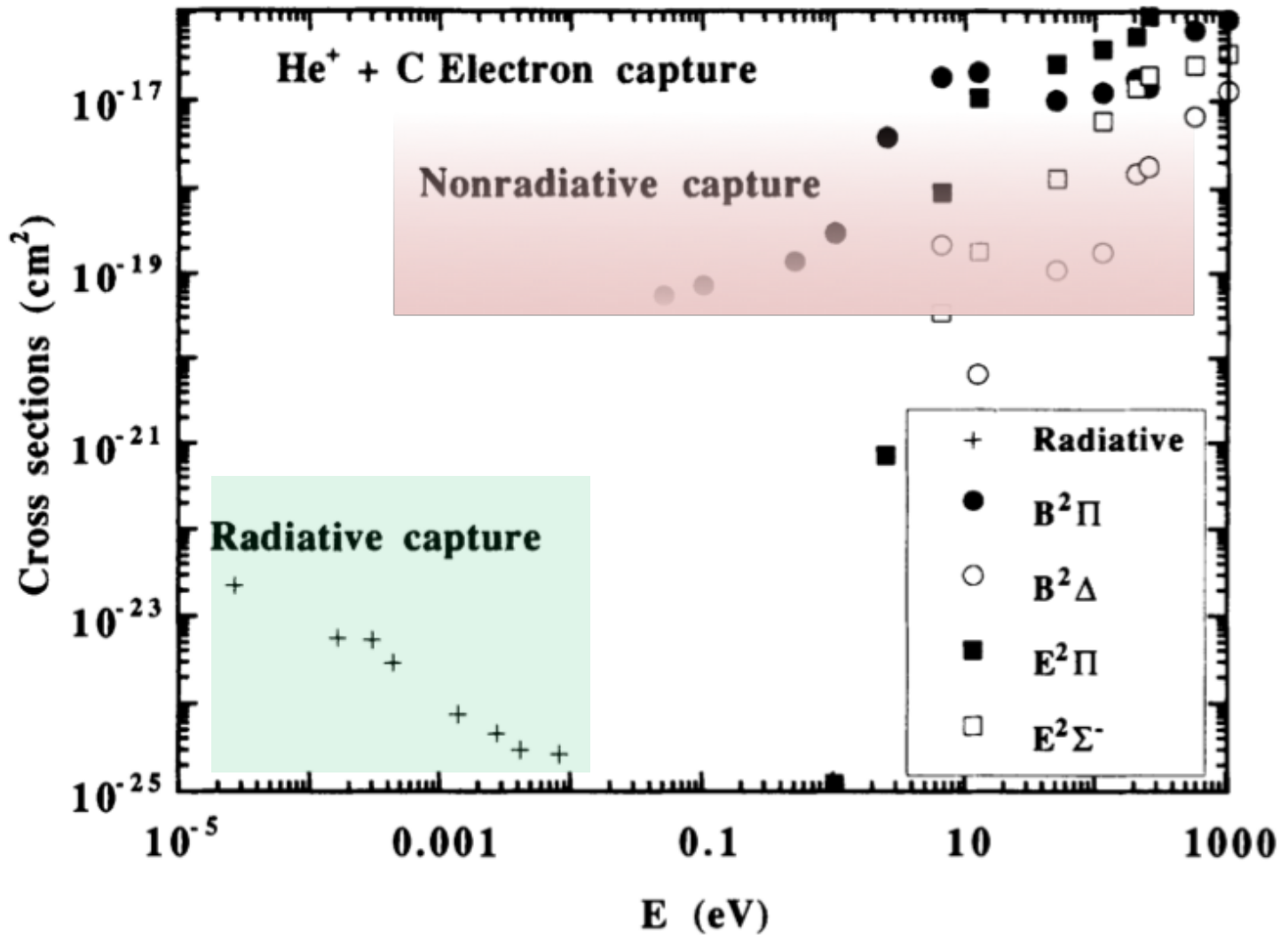
transfer:

Kimura, *et al.*

PRA

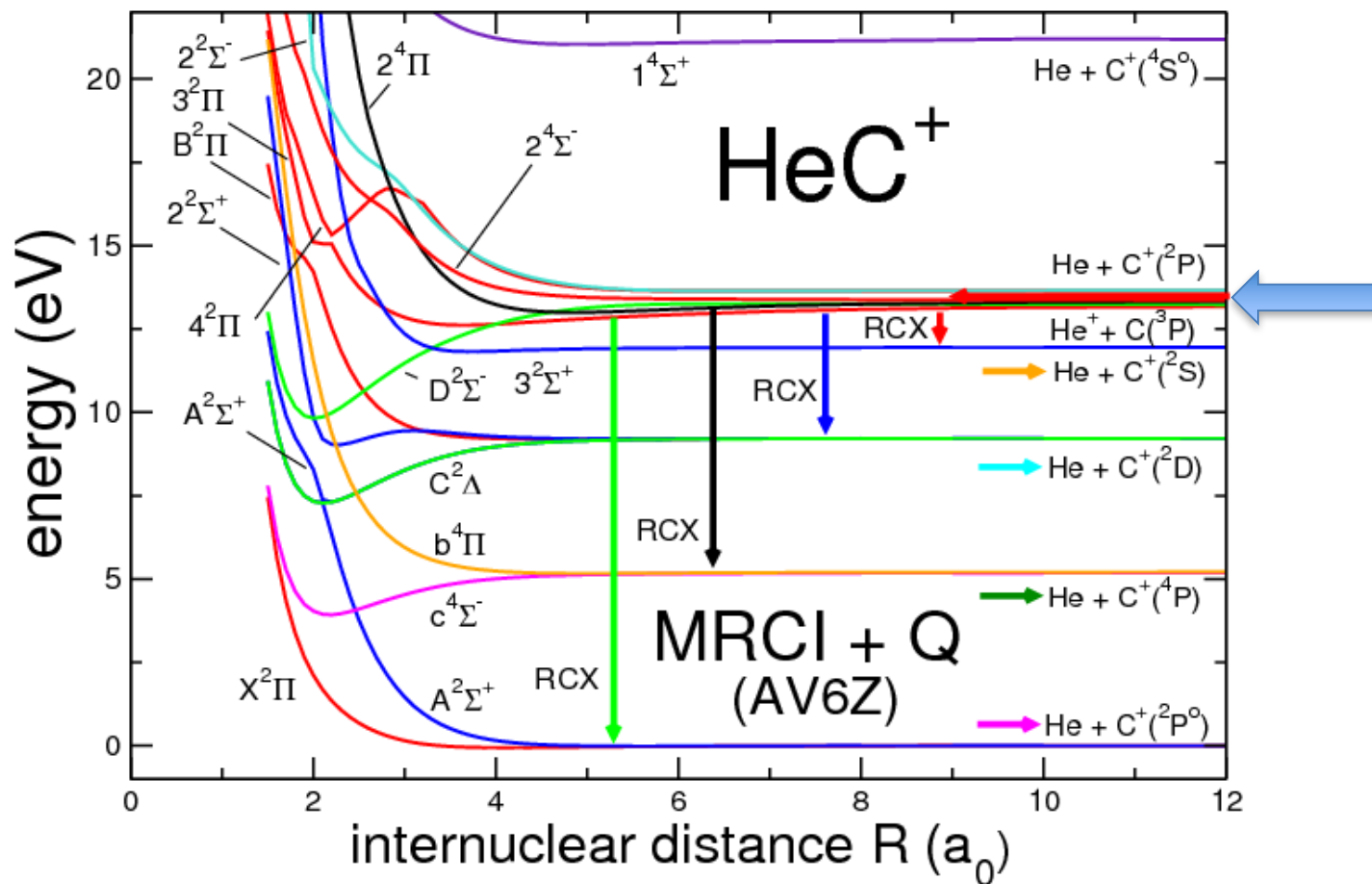
49, 2541

(1994)



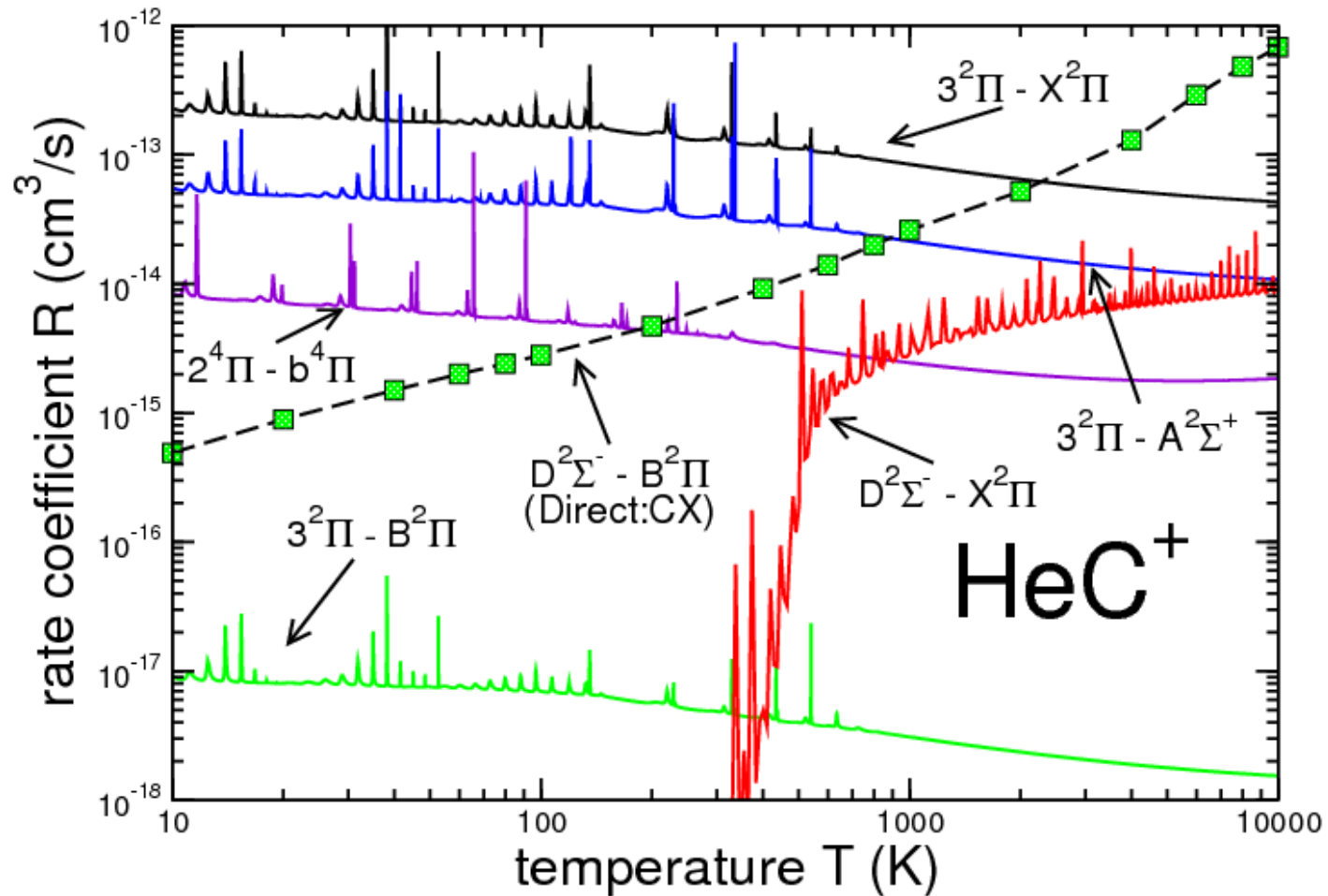
C-He⁺ energies

MOLPRO

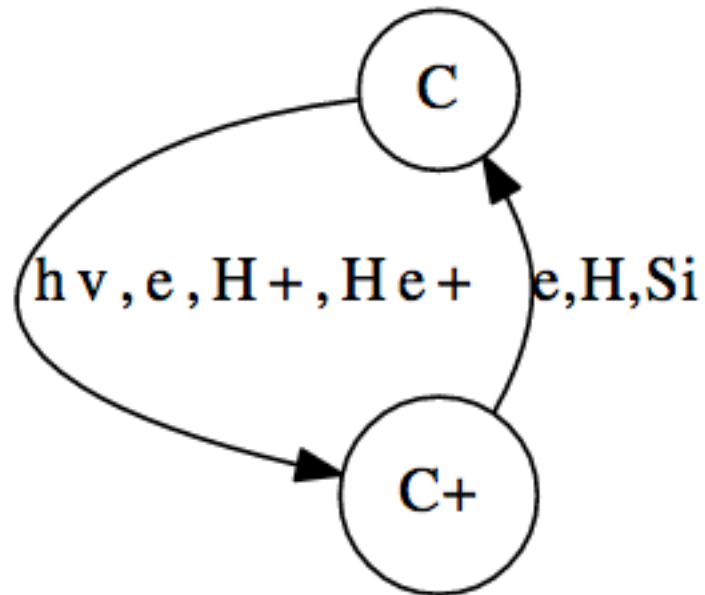


Babb & McLaughlin (2016)

C-He⁺ rate coefficients

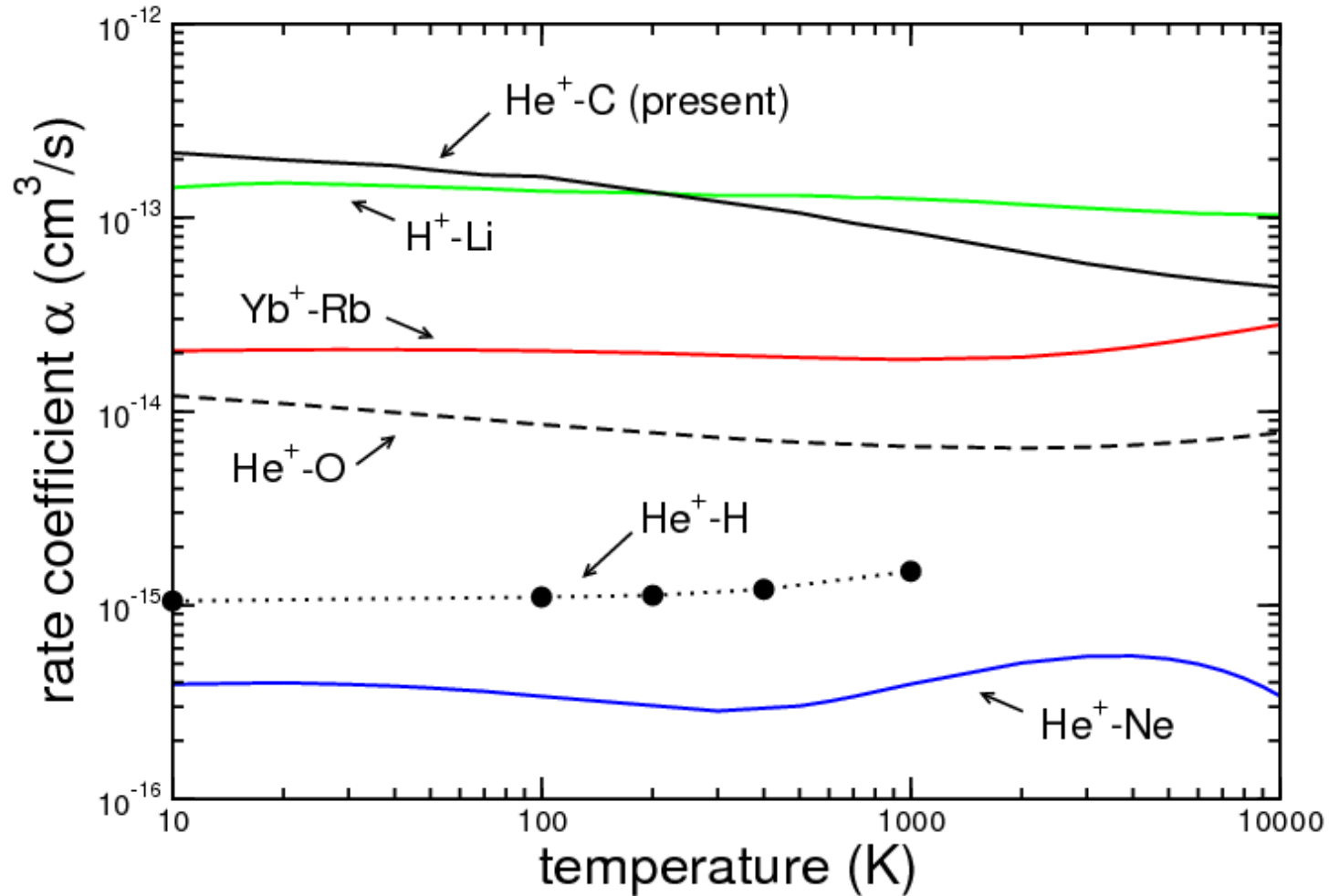


C-He⁺ network



Bovino et al. A&A 590 (2016)

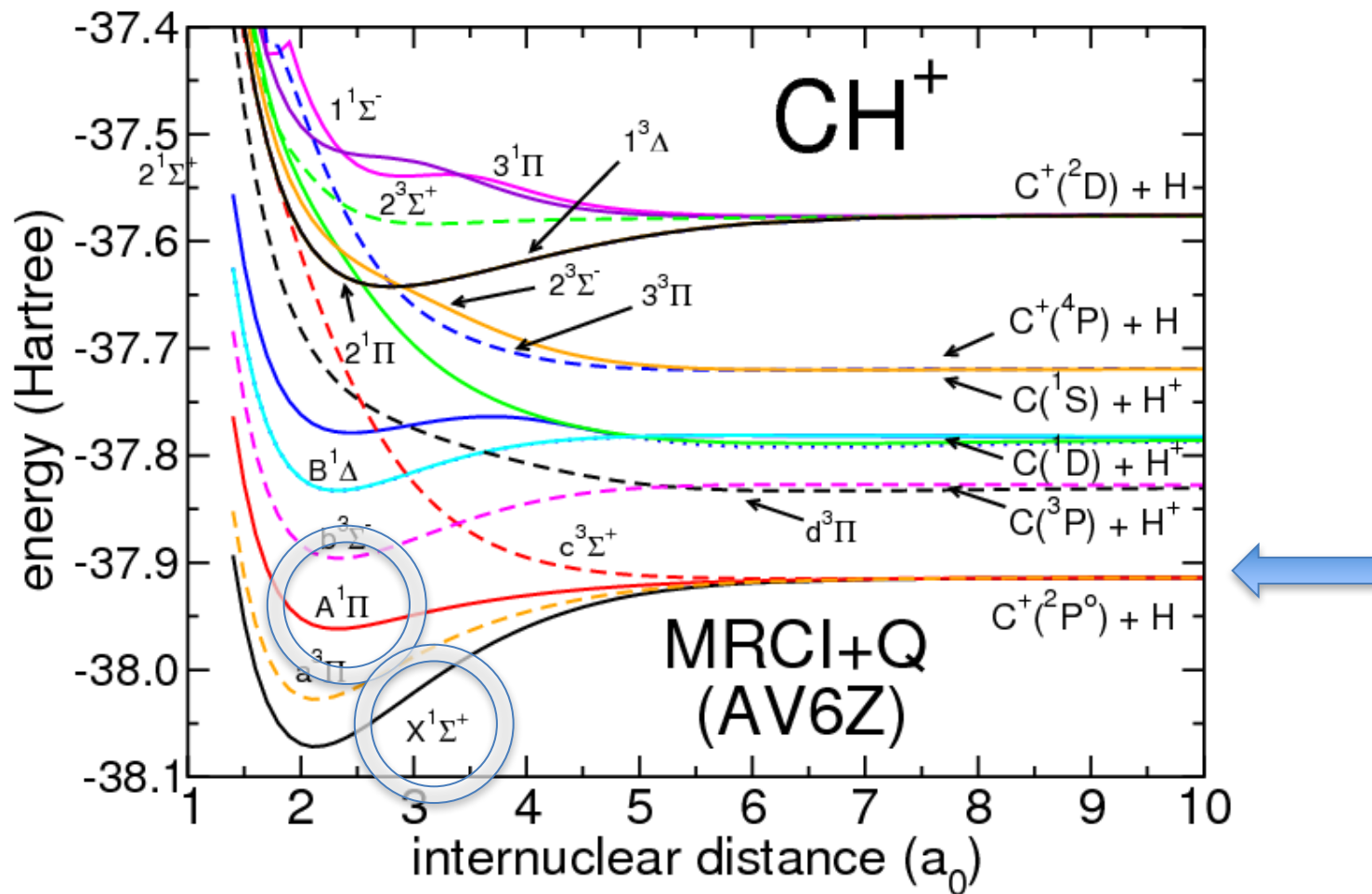
Comparisons: radiative charge transfer



CH⁺ prior work radiative association

Many papers over decades, explored for formation of CH⁺.

Now other mechanisms/reactions. All singlet transitions (A→X).



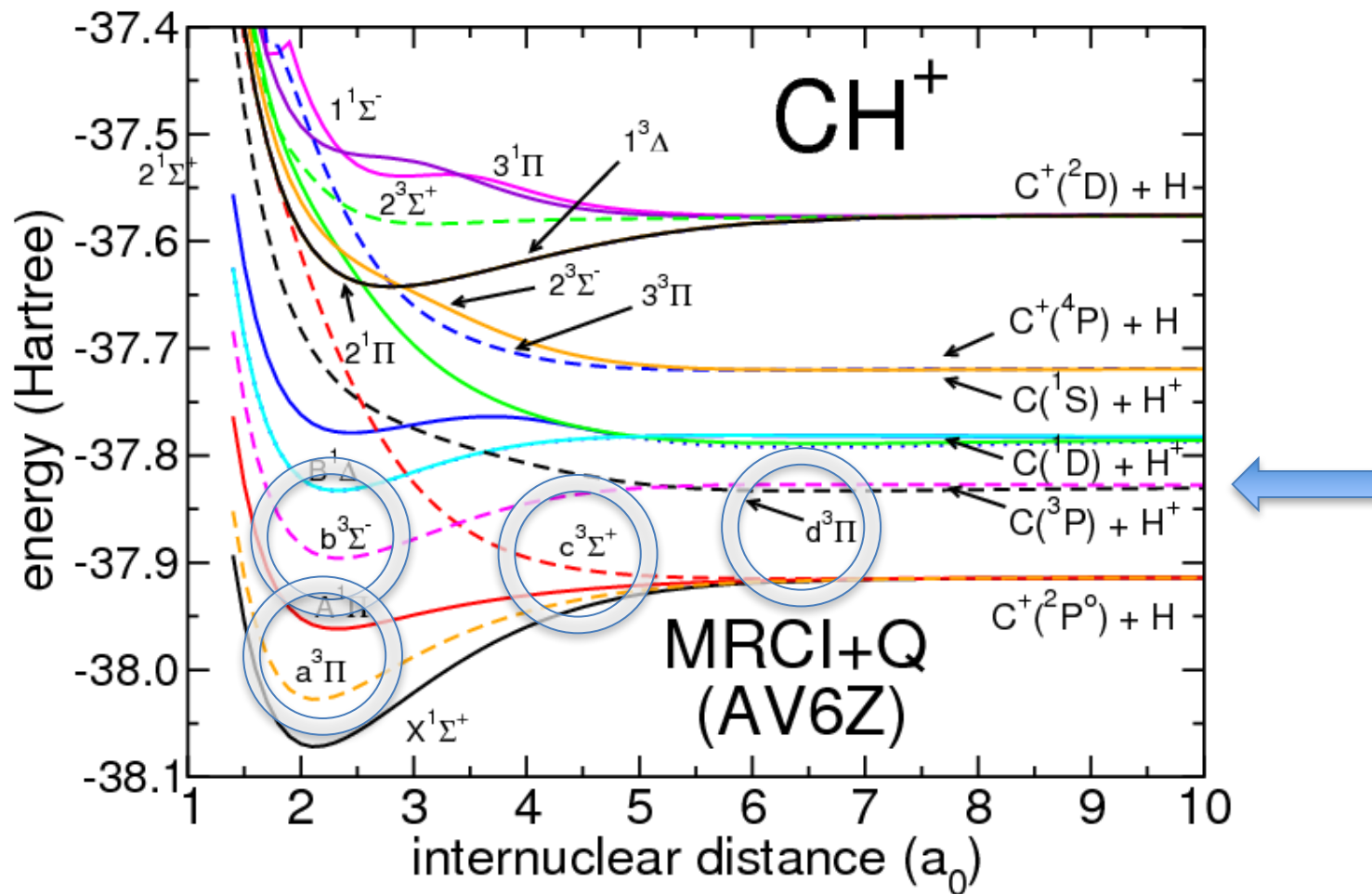
Radiative association cross sections

$$\sigma(E) \sim g \sum_J \sum_v \sigma_{vJ}(E)$$

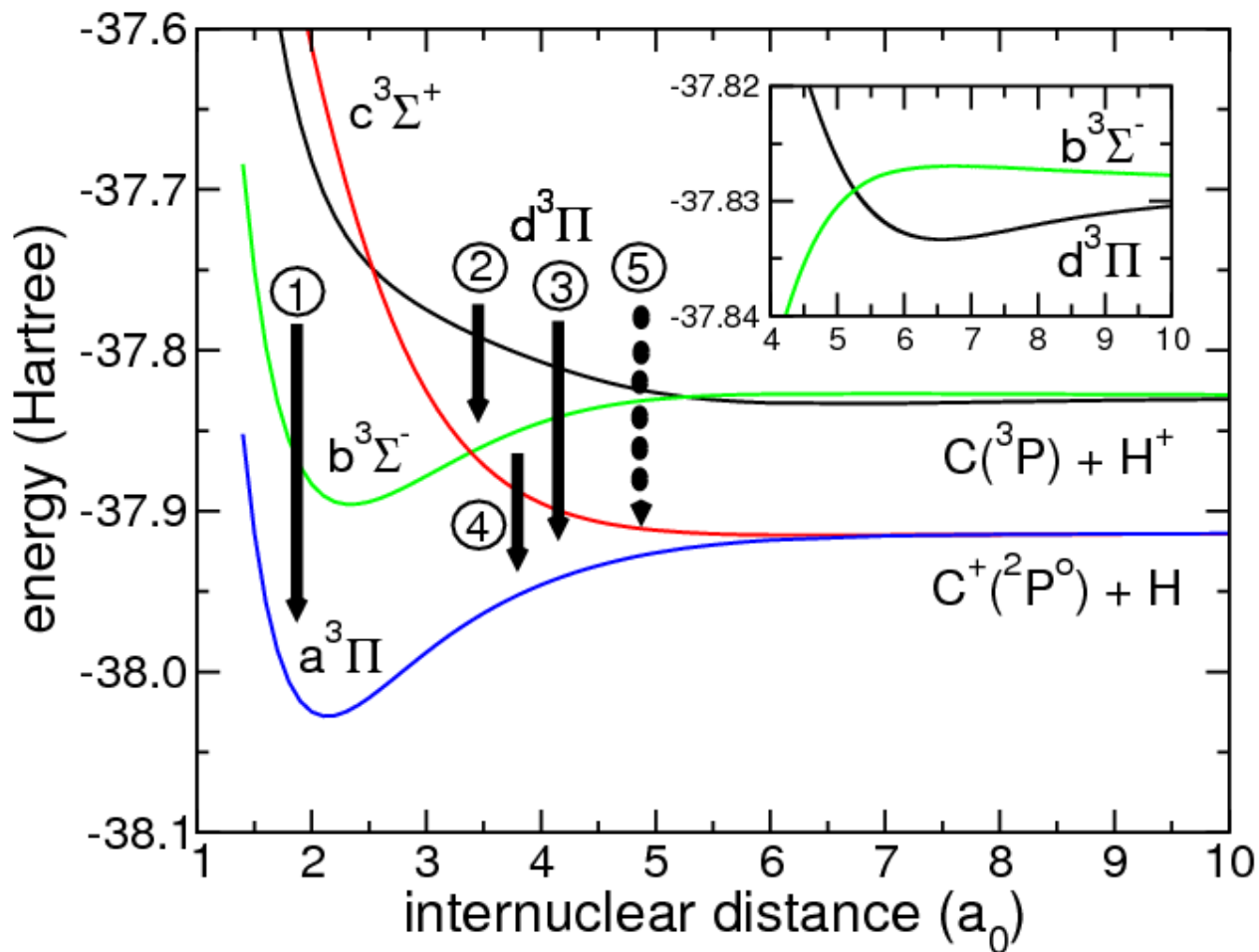
$$\sigma_{vJ}(E) \sim \frac{(h\nu)^3}{E} |\langle vJ | D(R) | k \rangle|^2$$

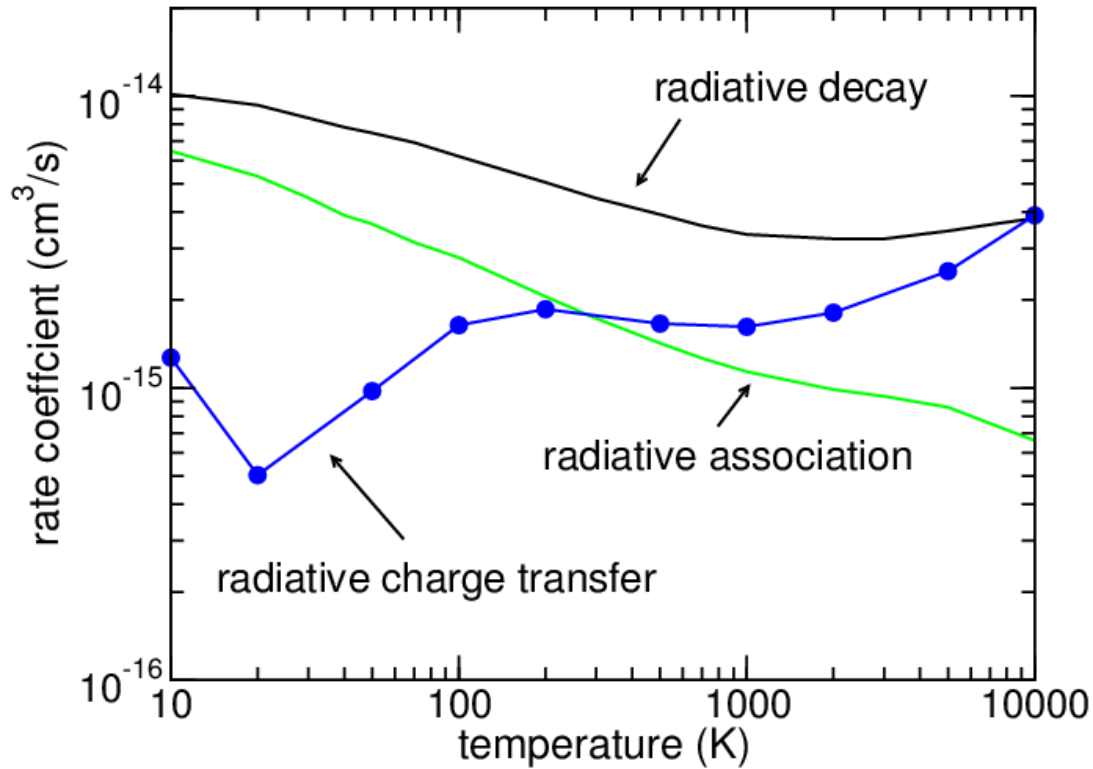
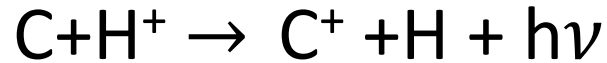
CH⁺ prior work radiative association

Consider, triplet transitions (d → a), (d → b), (d → c), (c → a), (b → a)



C-H⁺ radiative association / radiative charge transfer





$\text{C} + \text{H}^+ \rightarrow \text{CH}^+ + h\nu$
rate coefficient about 100× that for
 $\text{C}^+ + \text{H} \rightarrow \text{CH}^+ + h\nu$
... though need "C"

SiO - prior work radiative association

Found in Supernova 1987a “ejecta” (as was CO)

Calculations: $\text{Si} + \text{O} \rightarrow \text{SiO} + \text{photon}$

Andreazza *et al.* ApJ (1995)

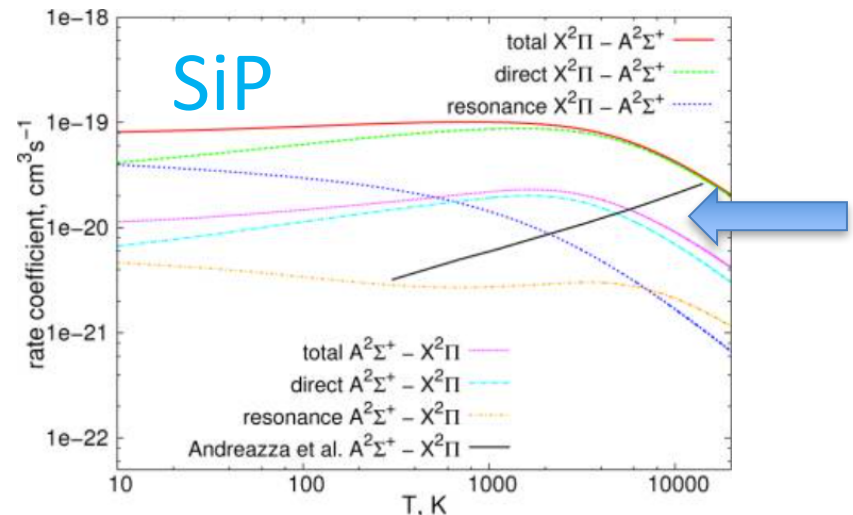
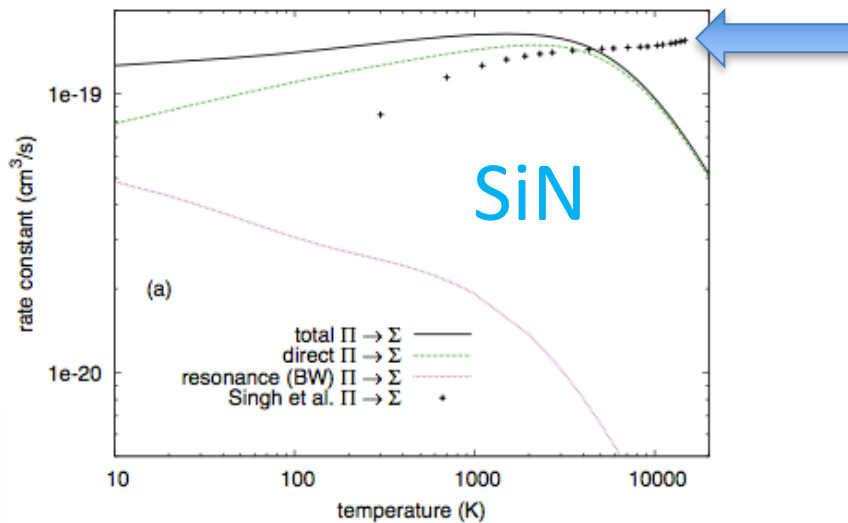
Problems?

Potential energy data

Inconsistencies vs. Andreazza+ ApJ (2006)

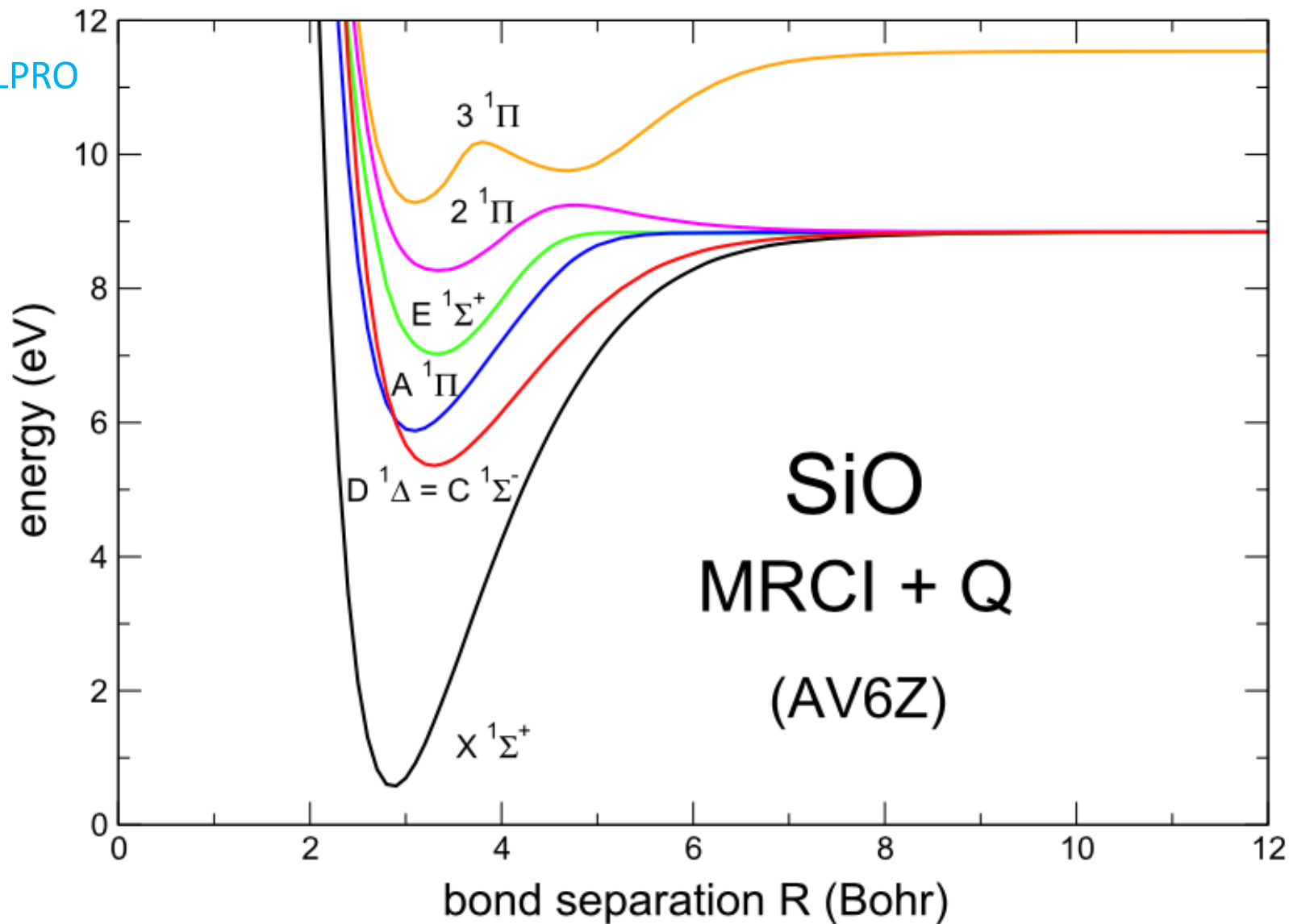
SiN — Gustafsson *et al.* JCP (2012)

SiP — Golubev *et al.* JPCA (2013)



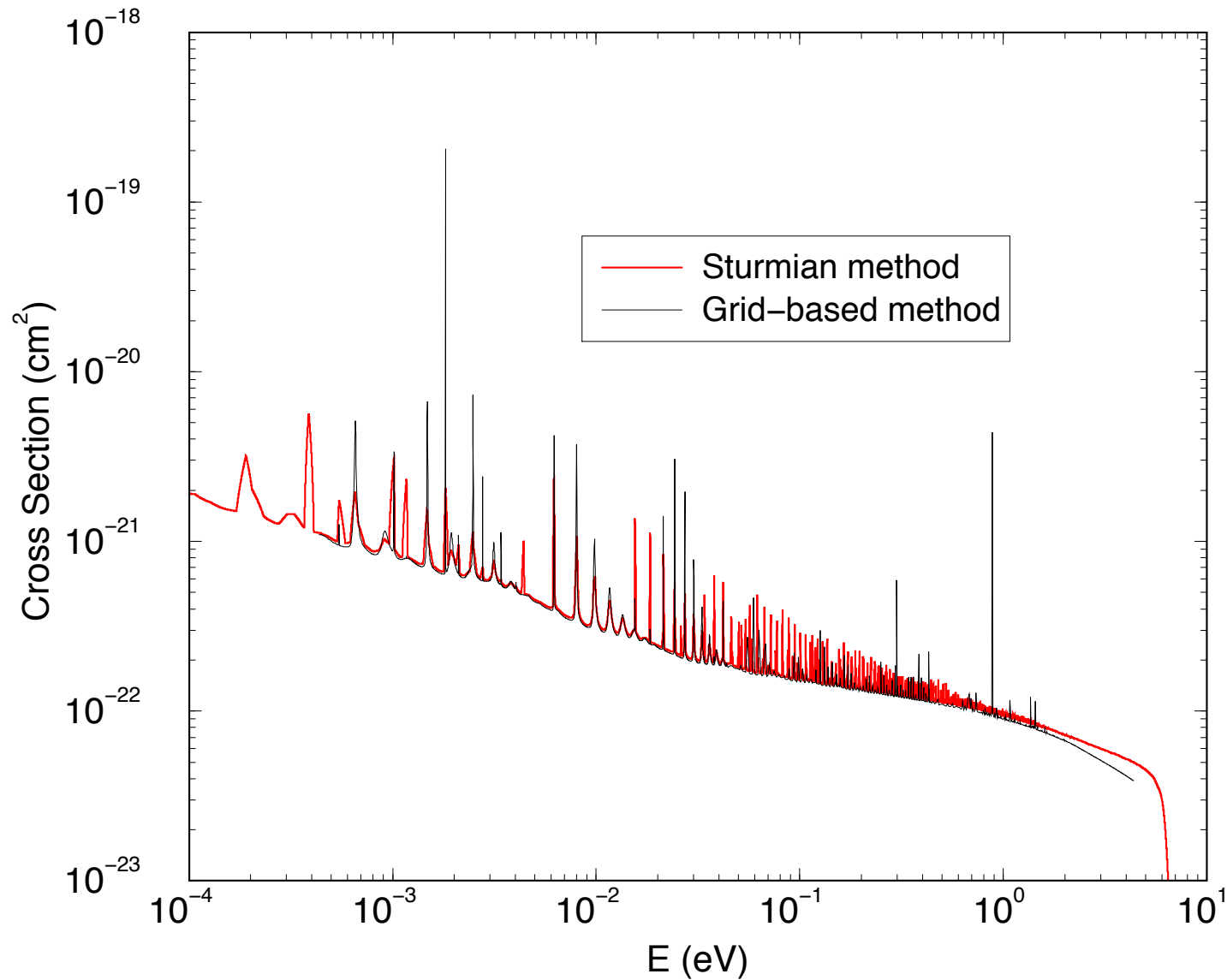
Potential energy data

MOLPRO



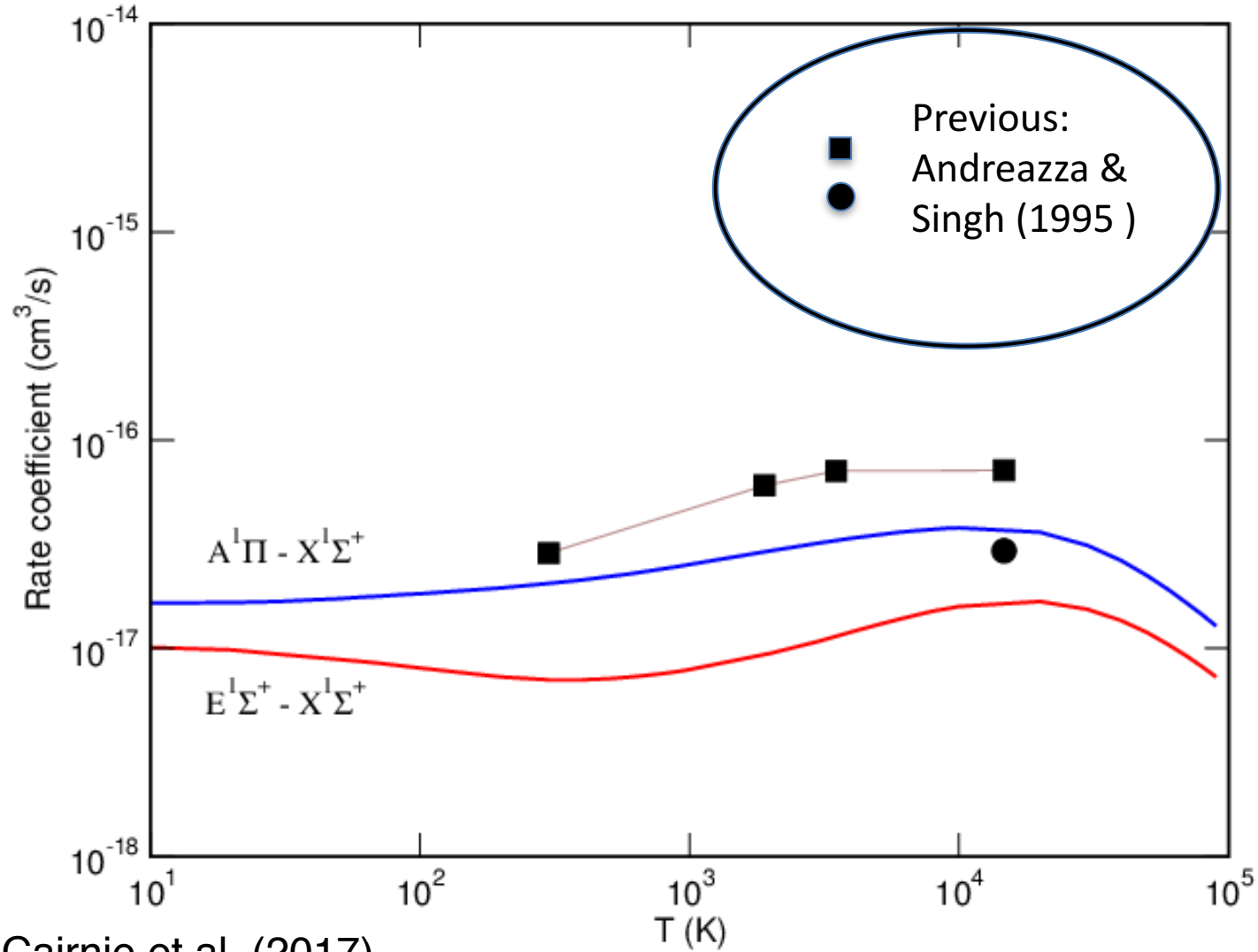
Forrey et al. (2016)

Cross sections: Two methods



Forrey et al. (2016); Cairnie et al. (2017)

Rate coefficients



Cairnie et al. (2017)

Postdoctoral fellowship in theoretical astrochemistry.

Research areas include, but are not limited to, primordial chemistry, and atomic and molecular processes in interstellar space and exoplanet atmospheres.

Applications deadline November 17, 2017.

Interested applicants are encouraged to visit itamp.harvard.edu for further information and access the online application process.

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ITAMP (NSF)



If you have ideas for workshops or visits ...
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